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## Inferring SRS character from time-dependent measurements of backscatter in NIF hohlraums

J. D. Moody, P. Michel, L. Divol, S. H. Glenzer, R. Town, E. A. Williams, S. Dixit, B. J. Macgowan, R. London, R. L. Berger, D. J. Strozzi, A. M. Rubenchik, K. N. Lafortune, R. K. Kirkwood, P. Datte, R. Hibbard, K. Krauter, E. Bond, L. Suter

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Snowmass Village, CO, United States  
June 13, 2010 through June 18, 2010

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# Inferring SRS character from time-dependent measurements of backscatter in NIF hohlraums

**Presentation to the  
40<sup>th</sup> Annual Anomalous Absorption Conference  
Snowmass Village, CO**

**J. D. Moody,** P. Michel, L. Divol, S. H. Glenzer, R. Town, E. A. Williams, S. Dixit, B. J. MacGowan, R. London, R. L. Berger,  
D. J. Strozzi, A. M. Rubenchik, K. N. LaFortune, R. K. Kirkwood, P. Datte, R. Hibbard, K. Krauter, E. Bond, and L. Suter  
*Lawrence Livermore National Laboratory, Livermore, CA 94550*

**14 June 2010**

**Lawrence Livermore National Laboratory**

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# Collaborators

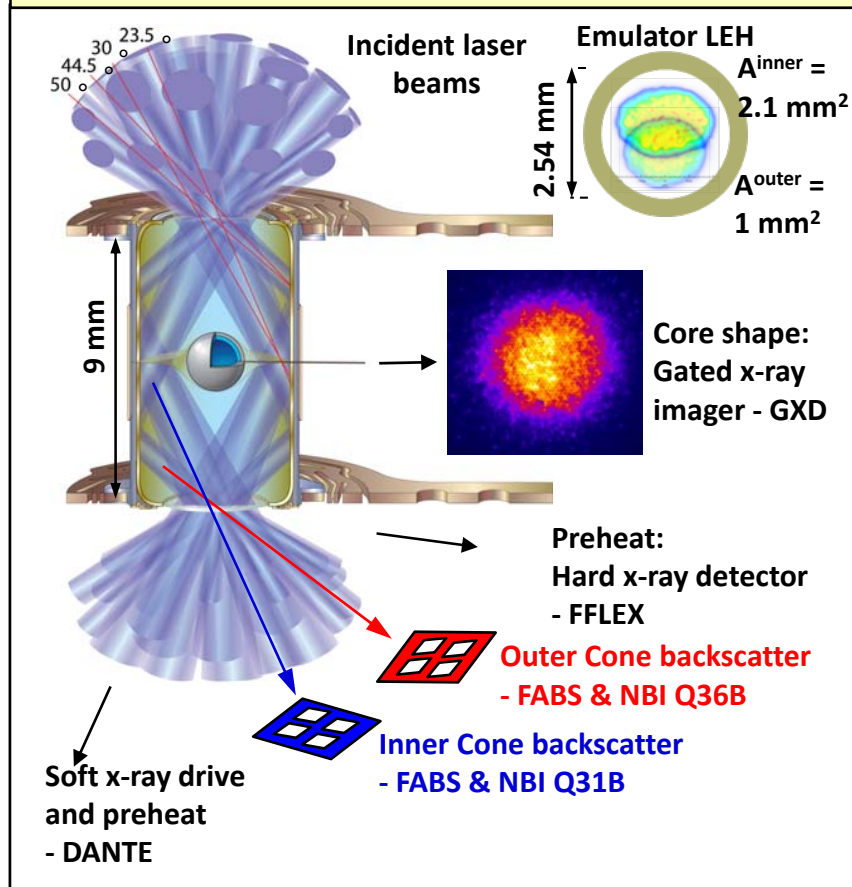
## **This work is done in collaboration with:**

*S. H. Glenzer, P. A. Michel, L. Divol, E. Bond, K. Knittel, C. Niemann[1], N. Meezan, B. J. MacGowan, B. Young, P. Datte, K. Krauter, R. Hibbard, J. Kilkenny, R. Wallace, J. Jackson, C. Gibson, S. Langer, B. Haid, R. London, G. Ross, G. Frieders, B. Beeman, D. Pigg, J. Nelson, R. Shelton, J. Laney, J. Nelson, B. Golick, J. D. Kilkenny, D. Larson, J. Atherton, M. Bowers, E. Williams, D. Hinkel, J. Kline, S. Dixit, M. Jackson, G. Deis, R. Robinson, G. Vergel de Dios, J. Moody, L. Bertolini, T. Lee, M. Vitalich, S. Shiromizu, M. Richardson, T. Malsbury, B. Langdon, W. Hsing, S. Vernon, S. Azvedo, J. Kamperschroer, E. Ng, L. Bezerides, L. Belk, R. Beeler, A. Casey, R. Robinson, D. Martin, S. Andrews, M. Fischer, and L. Suter*

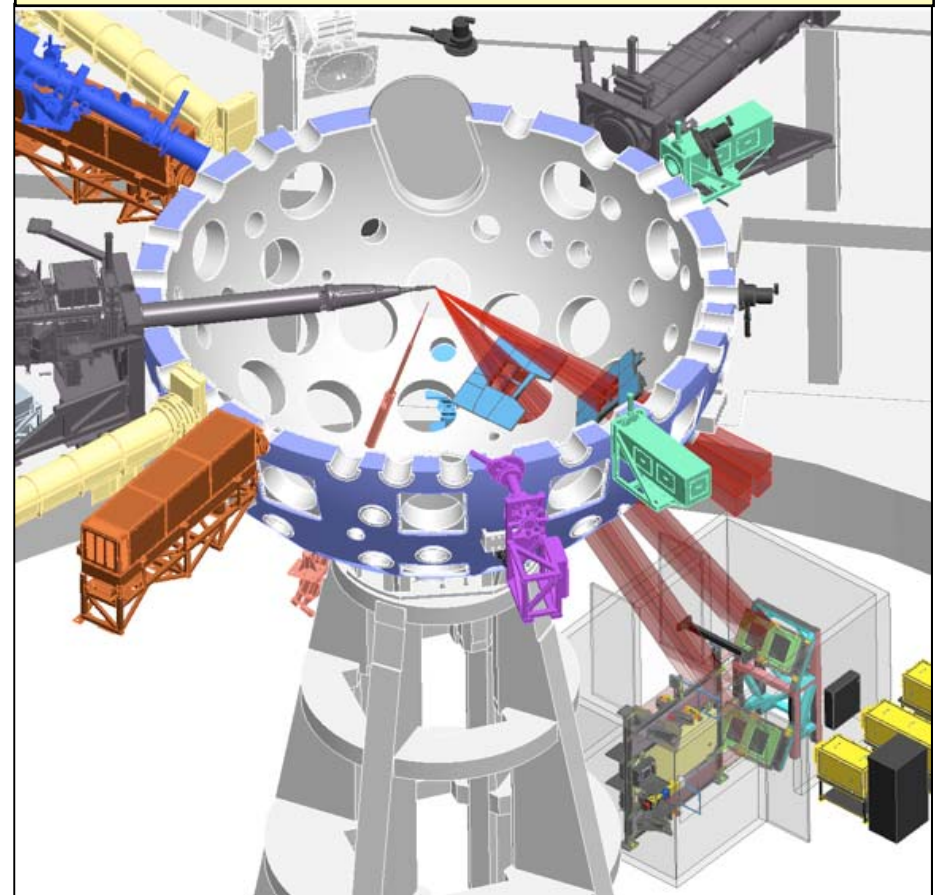
*[1] UCLA, Los Angeles, CA 90095-1594*

# The backscatter measurement is important for quantifying the energetics aspects of ignition hohlraums

Hohlraum heated by 192 beams  
emulating ignition conditions

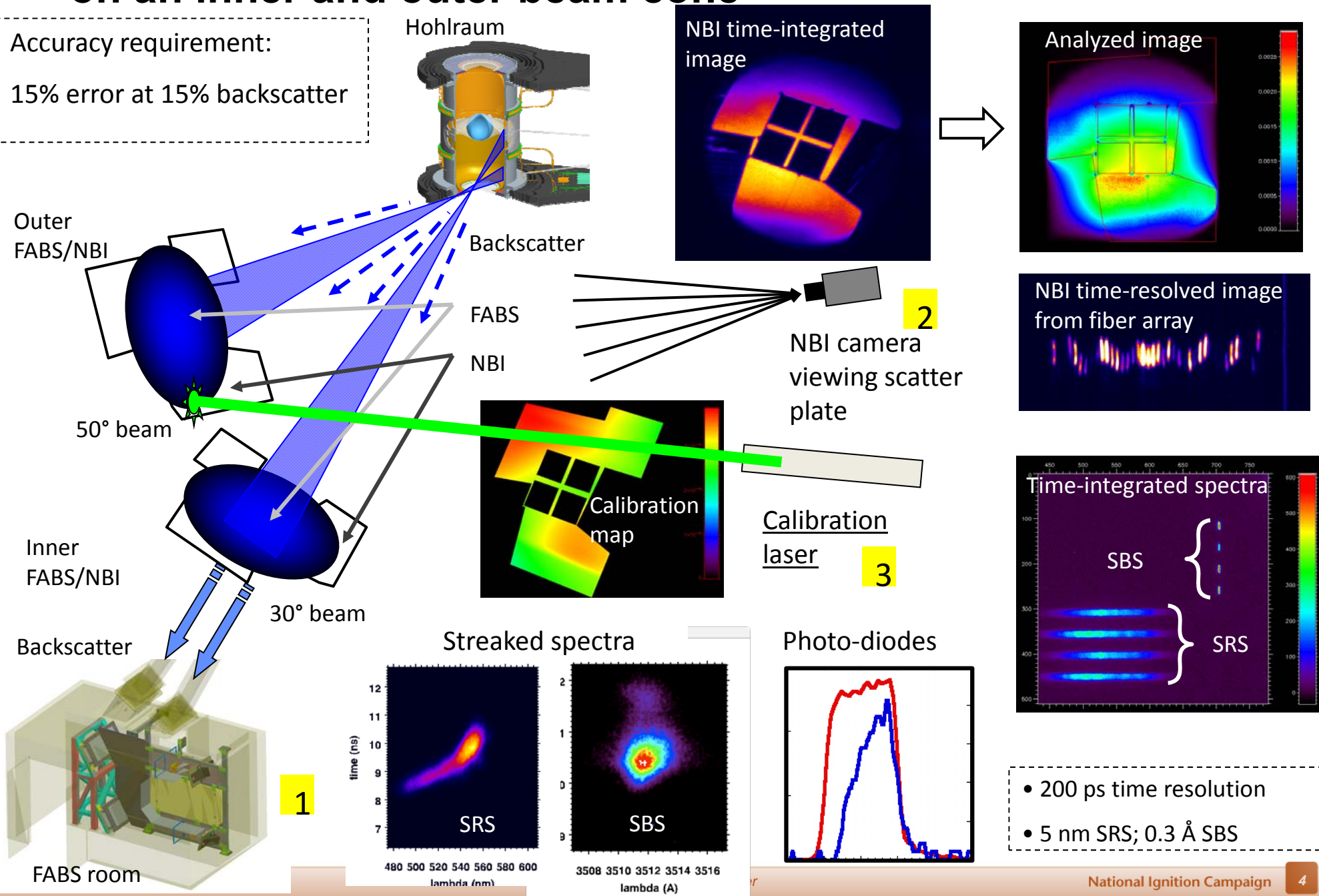


Backscatter instruments (FABS / NBI)  
are installed on two beam quads



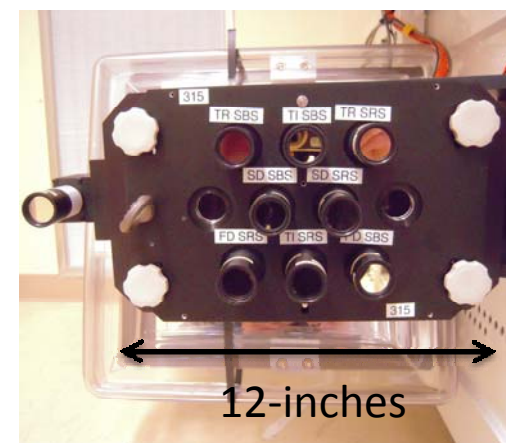
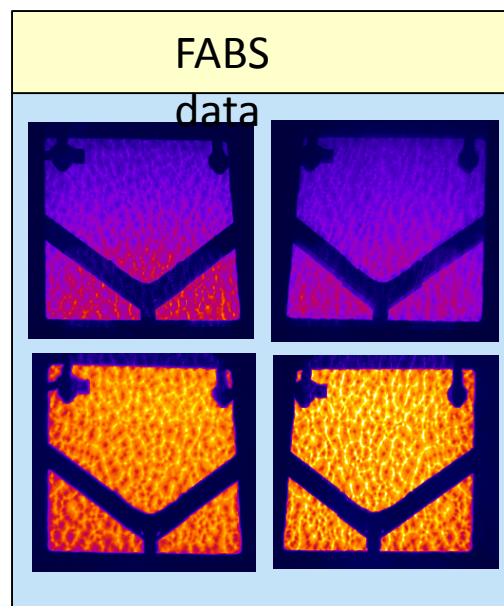
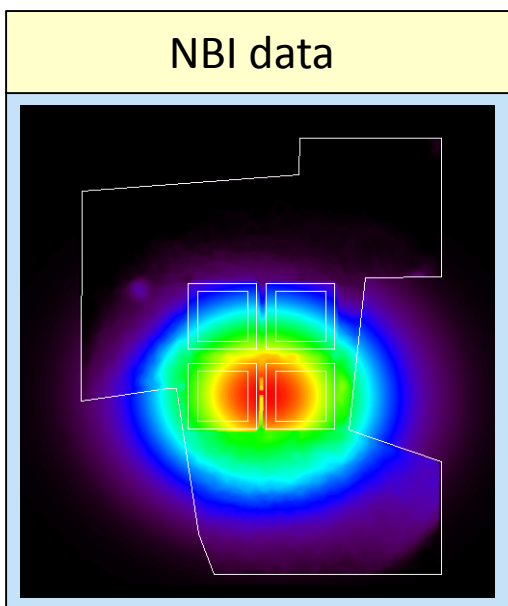
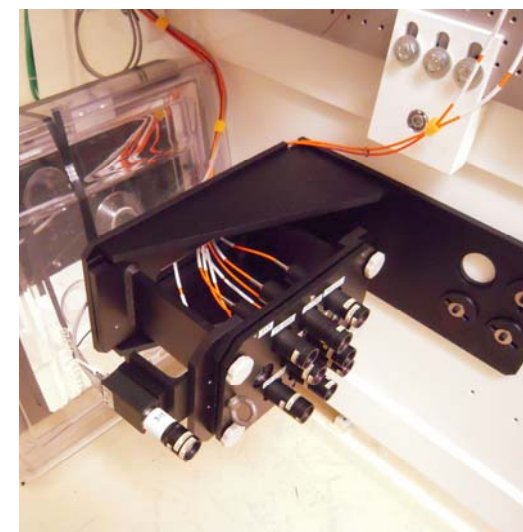
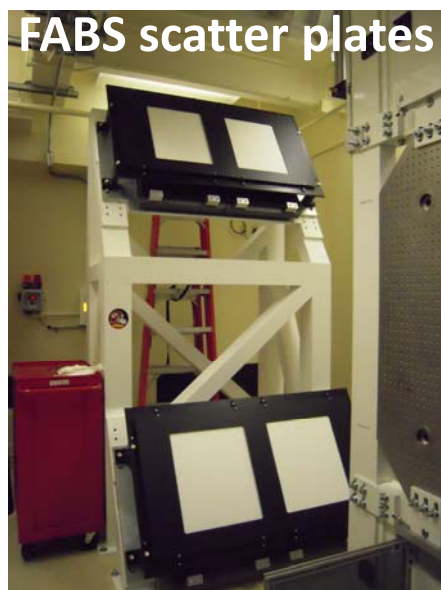
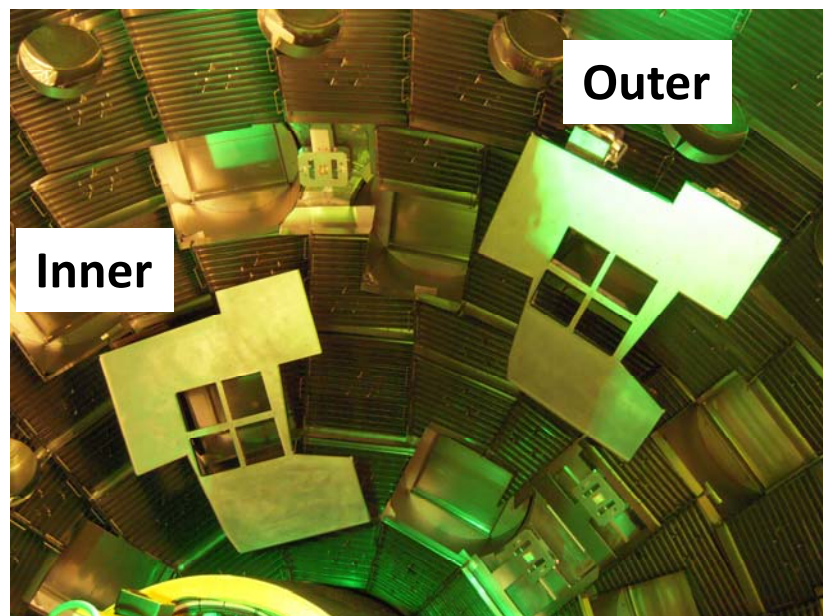
# We have installed backscatter diagnostics on an inner and outer beam cone

Accuracy requirement:  
15% error at 15% backscatter





# Backscattered instruments installed on NIF

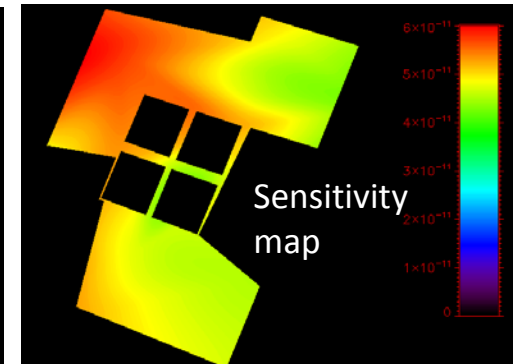
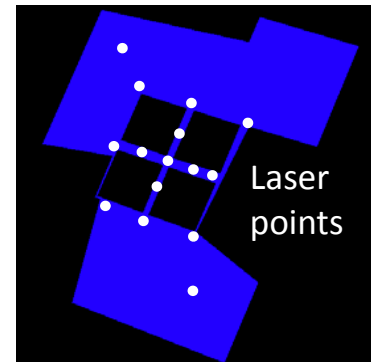


# Backscatter calibration combines several techniques to achieve $\geq 15\%$ relative accuracy

- 1) Pulsed laser calibration - 2 wavelengths, S and P polarization. Calibrates FABS diodes for *power* and NBI sensitivity



*Pulsed laser*

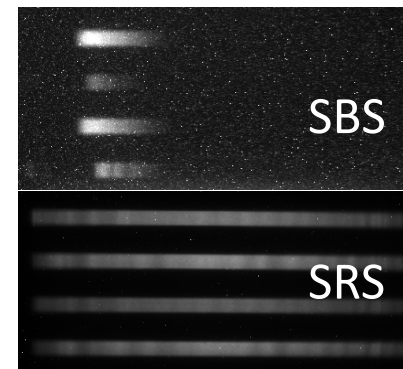


*14% error*

- 2) White light Xe lamp. Calibrates calibration spectrometer for FABS energy



*Xe lamp*



*20% error*

*17% error*

*Optical filters*



*Pencil calibration lamp*



*Wavelength*

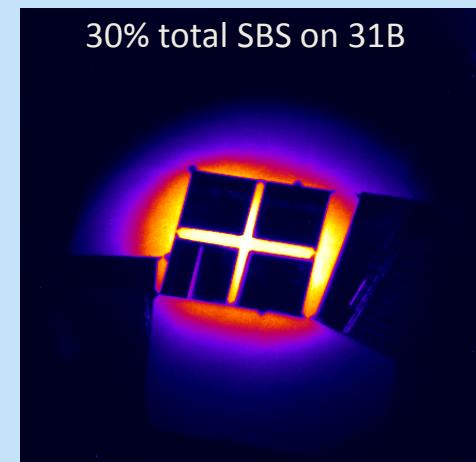
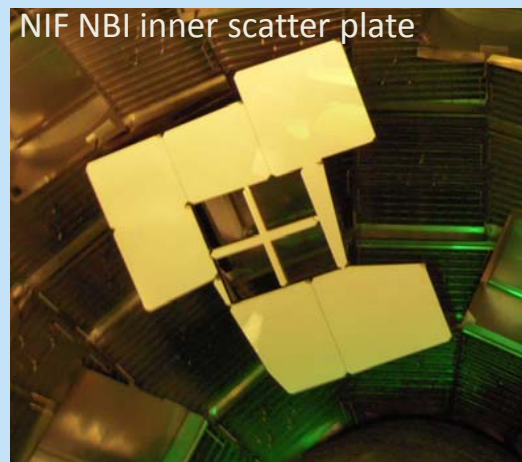
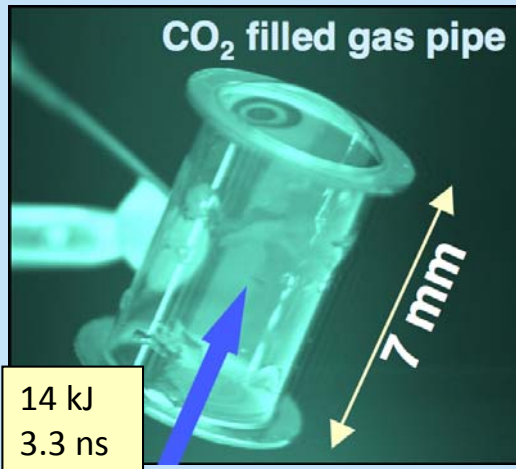
- 3) Spectrometers are calibrated using standard Hg or Ne calibration lamps. Instruments and filters are calibrated off-line

**FABS / NBI calibration is based on several independent measurements. Accuracy estimates are determined by the relative agreement between these**

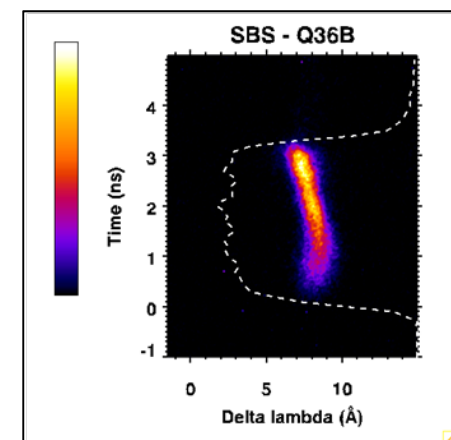
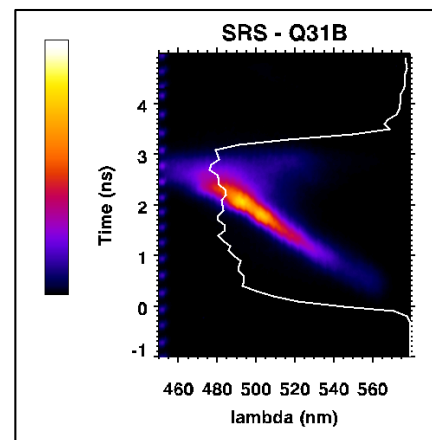
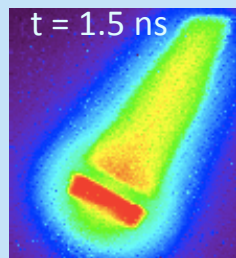
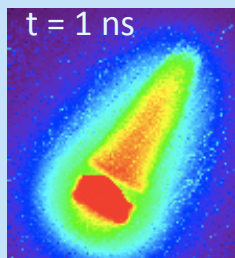


# Large scale length 7 mm-long plasmas have been produced with 1 quad of beams to activate scattering diagnostics

Near Backscatter Imagers measure SBS and SRS around inner (Quad 31B) and Outer (Quad 36B) beams



Gated x-ray data indicate burnthrough at  $t = 1.5$  ns



# We have performed experiments to measure the backscatter from cryogenic gas-filled hohlraums

Cryogenic hohlraum target with shields; He-H<sub>2</sub> gas-fill at 20.6 K

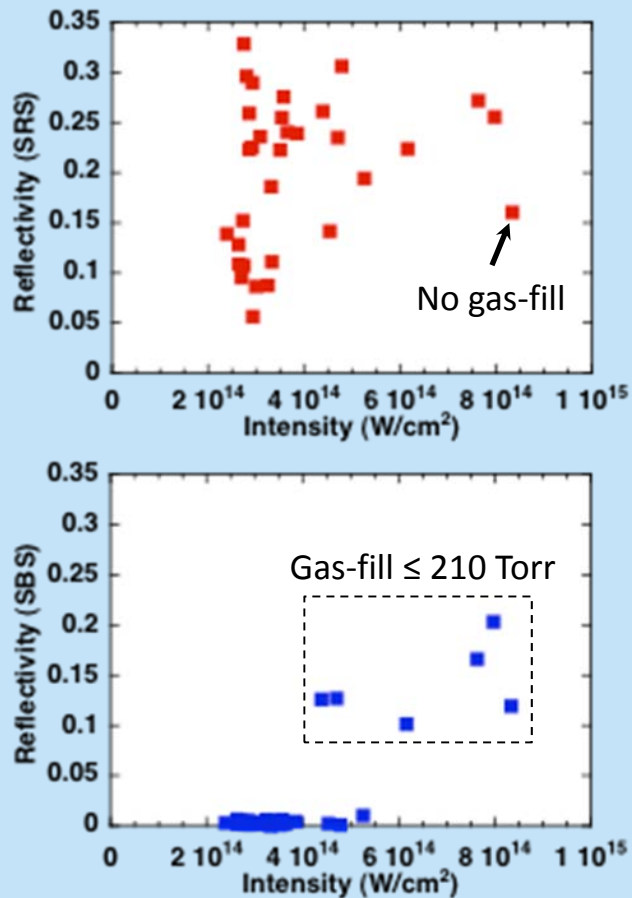


Some important results from hohlraum experiments:

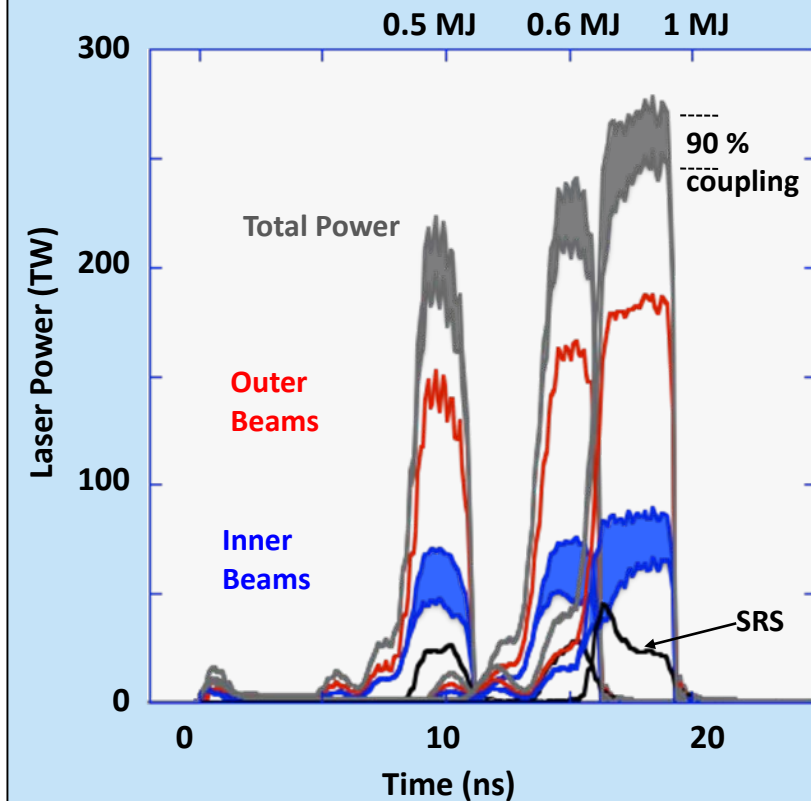
- LEH liners give high backscatter
- Pure He is better than He/H<sub>2</sub> fill
- Pure Au is as good as Au/B
- Checkerboard PS is better
- $\Delta\lambda$  tuning capability is important

# Inner beam backscatter is $\leq 10\%$

Inner SRS and SBS from hohlraums



Total SRS is  $\leq 10\%$  up to 1 MJ

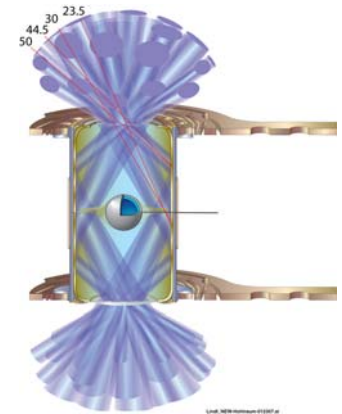
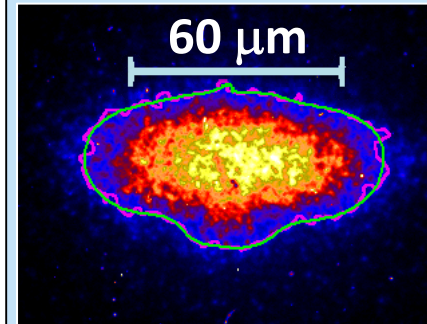
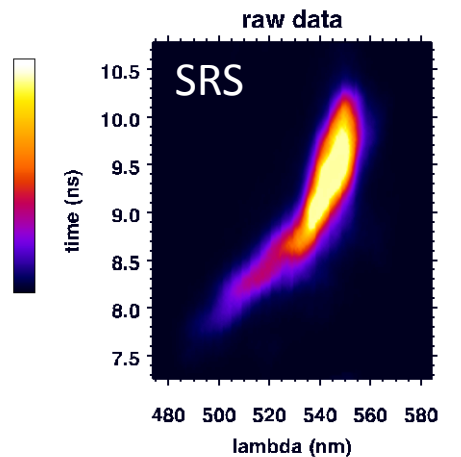


# Wavelength tuning counteracts the SRS inner-beam losses in cryogenic gas-filled hohlraums

Inner and outer cone wavelength separation:  $\Delta\lambda = 1.5 \text{ \AA}$

N090902 [12.27 kJ on 31B]

SRS = 23.6%  
He/H plasma

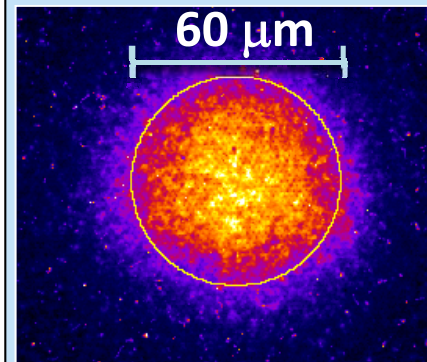
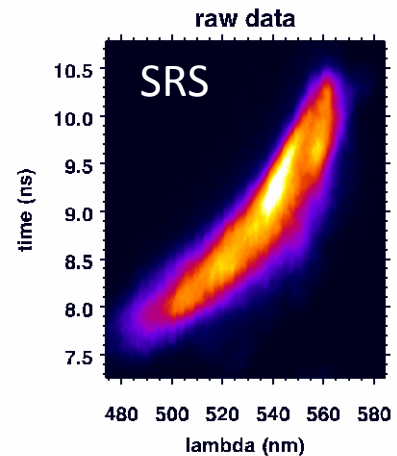


Hohlraum  
orientation

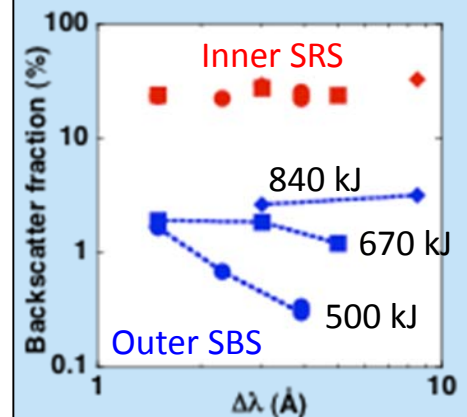
Inner and outer cone wavelength separation:  $\Delta\lambda = 3.9 \text{ \AA}$

N090905 [13.74 kJ on 31B]

SRS = 25.5%  
He/H plasma



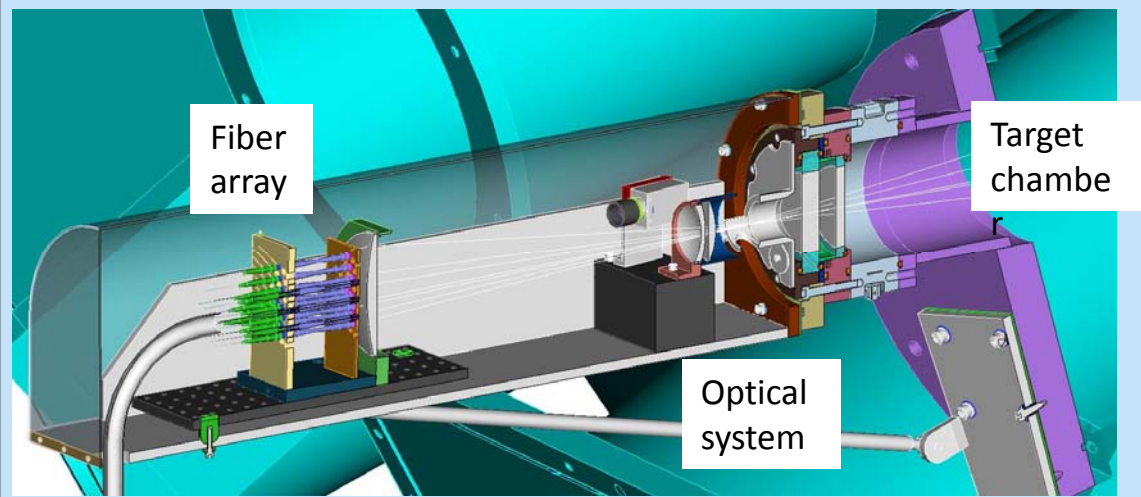
Outer SRS



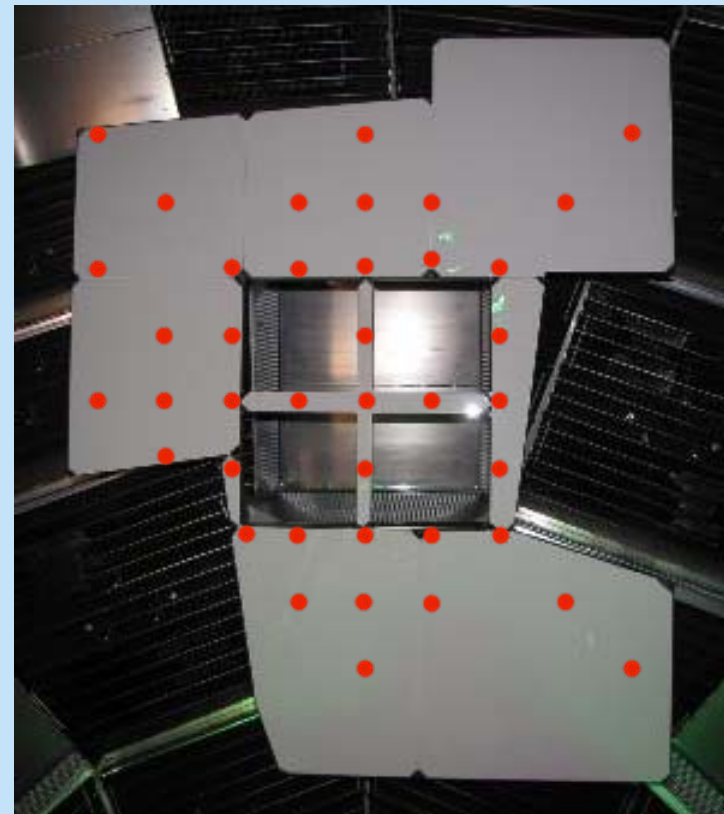


# Time-resolved NBI utilizes 40 fibers and a streak camera

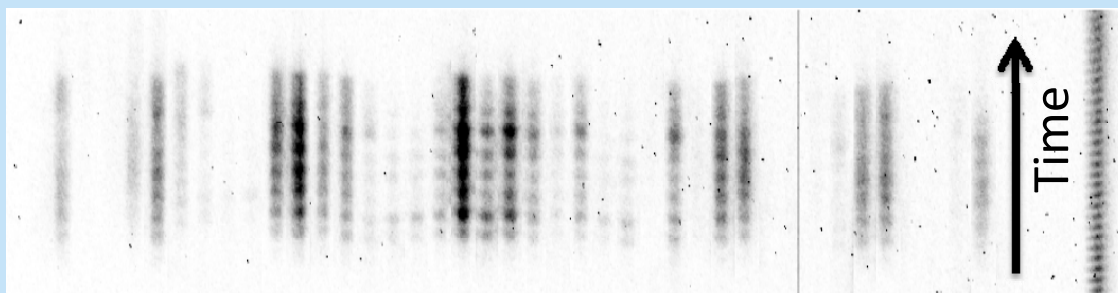
Sketch of NBI-time-resolved system



Scatter plate with fiber views

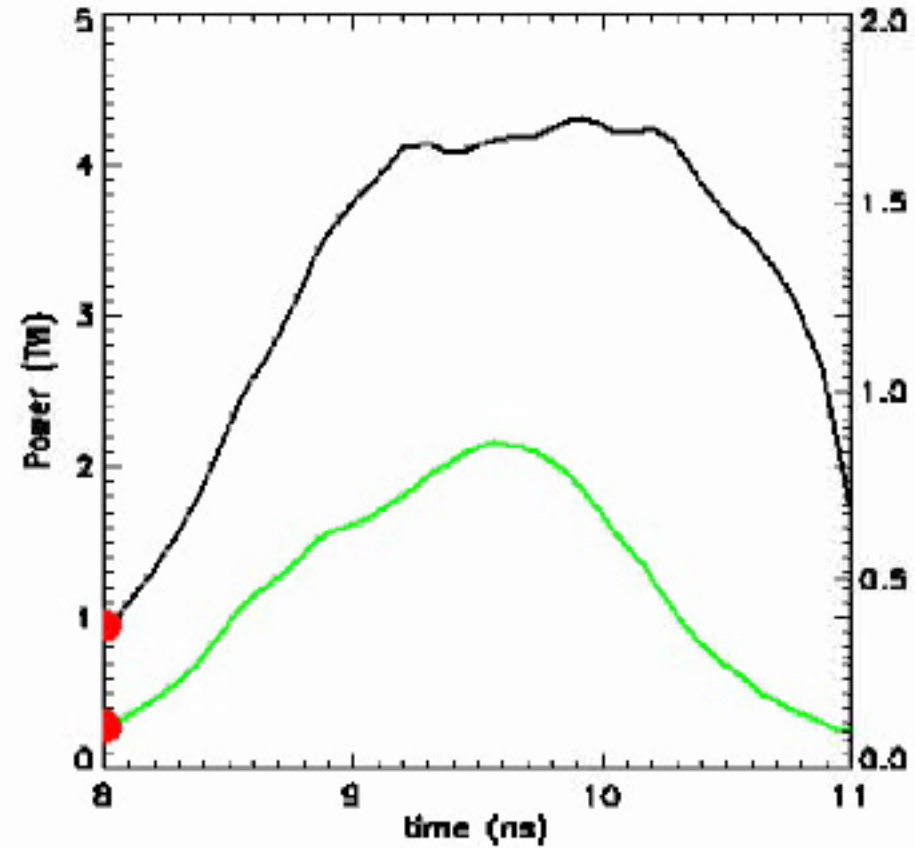
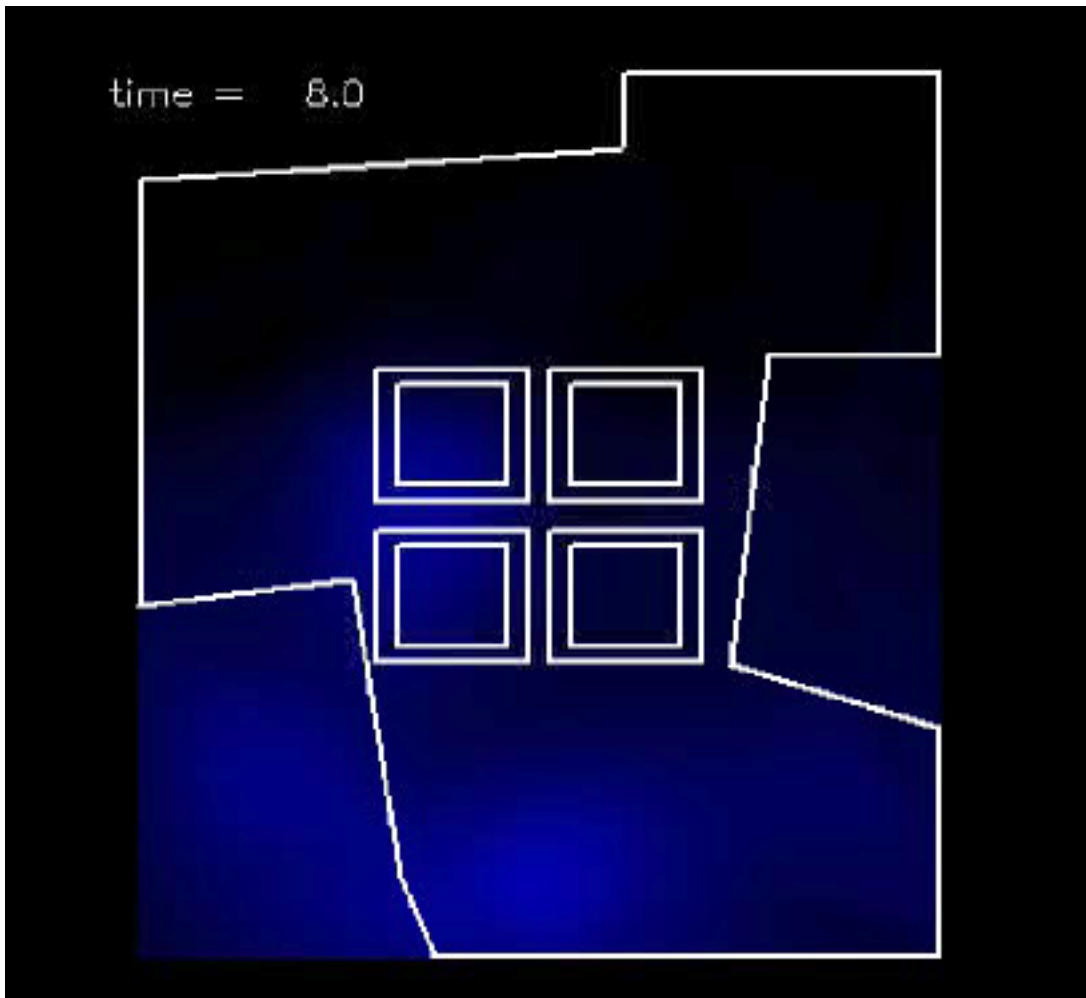


Fiber streaks



Data corrected for propagation time

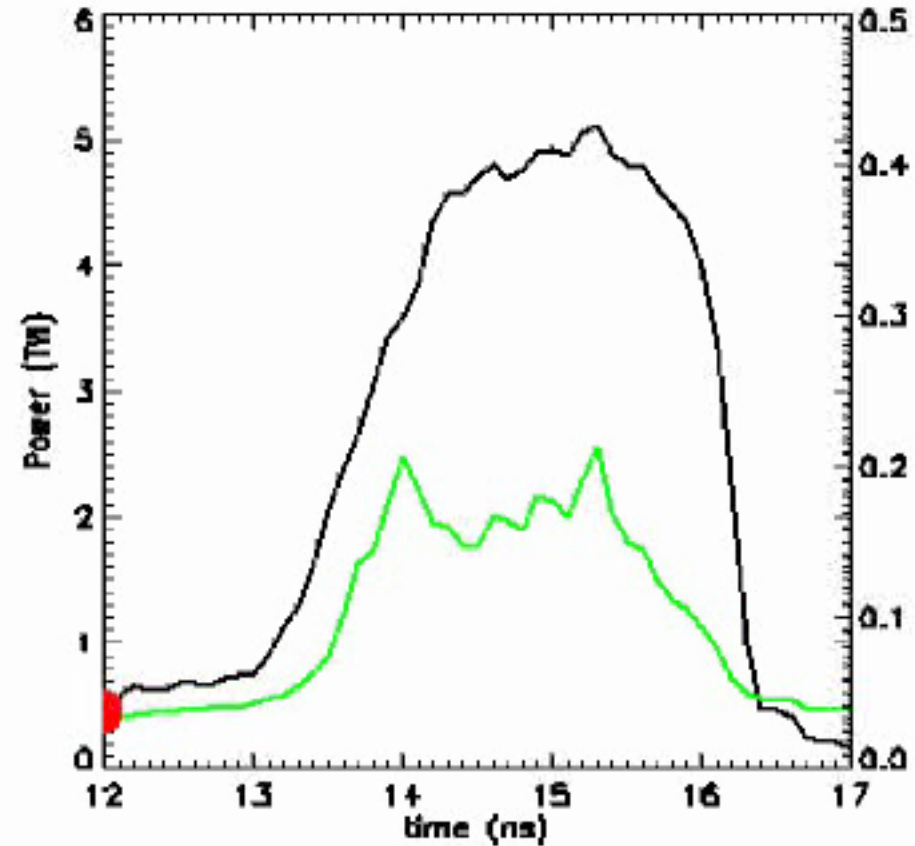
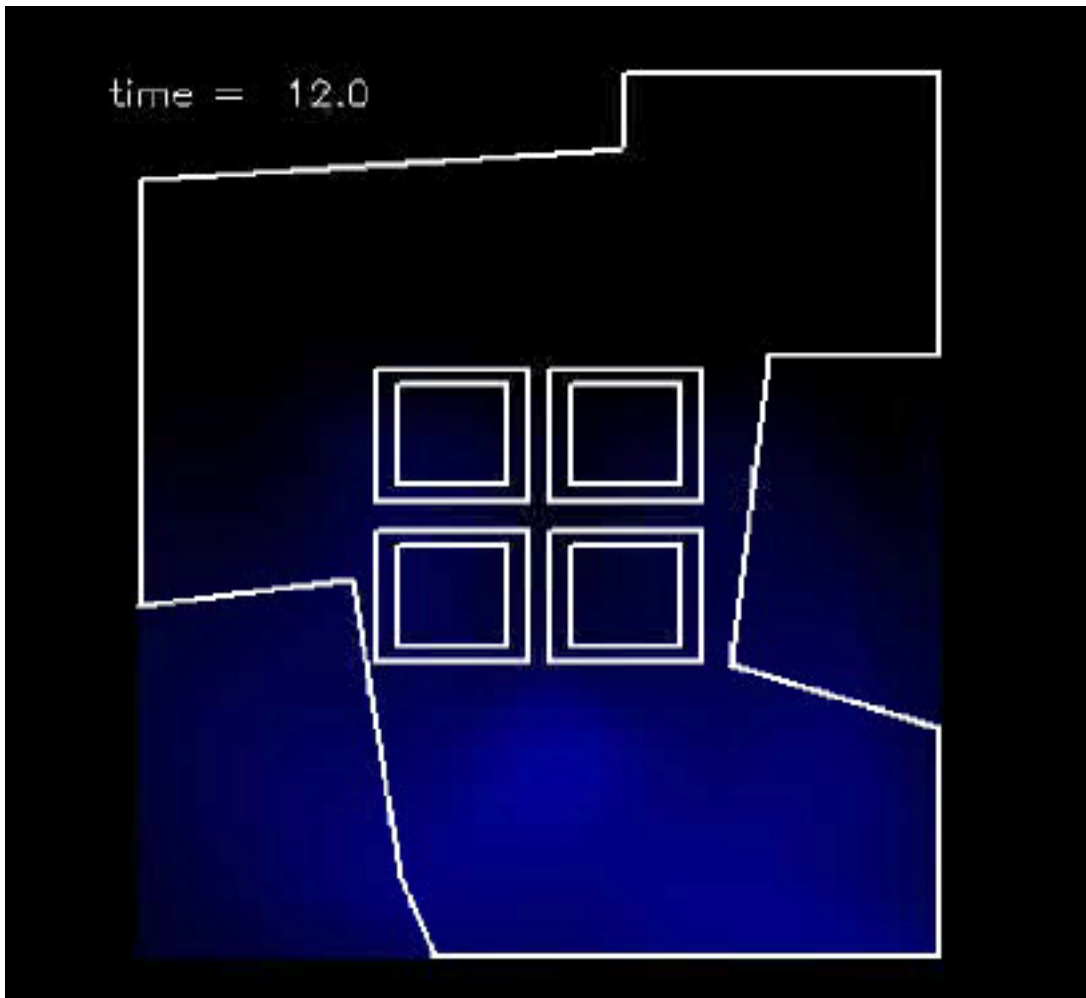
# NBI movie for a hohlraum with an LEH liner



Shot N0911107

29 % SRS

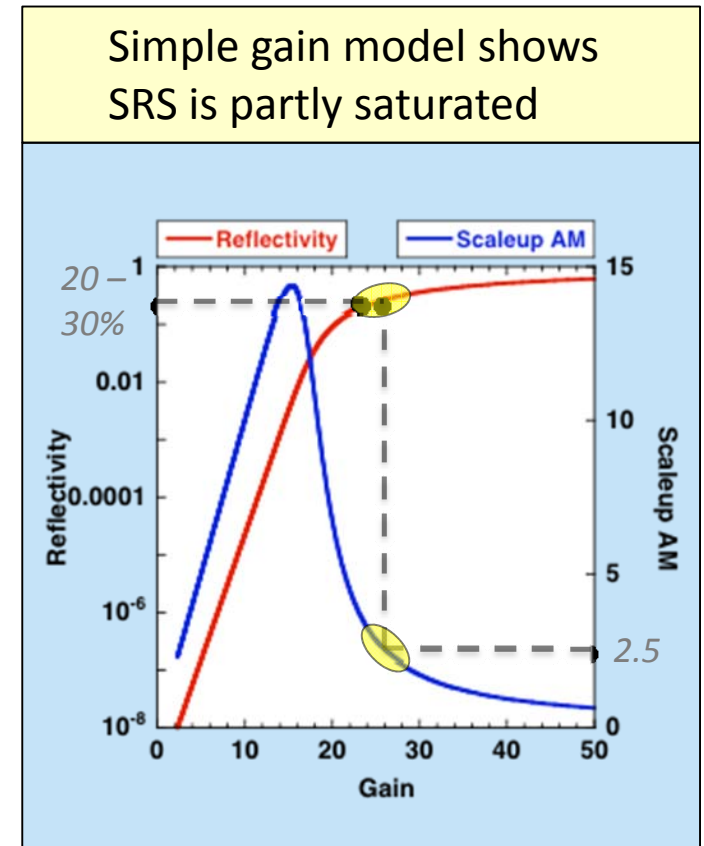
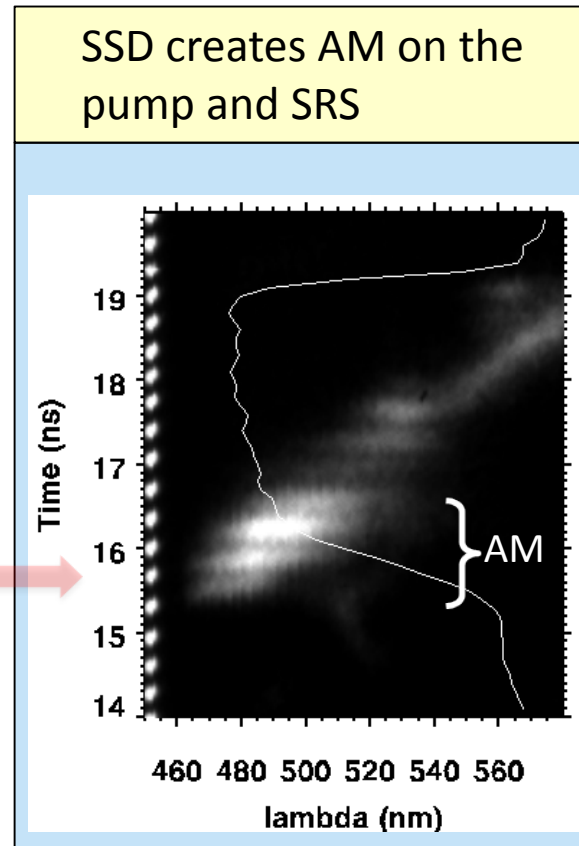
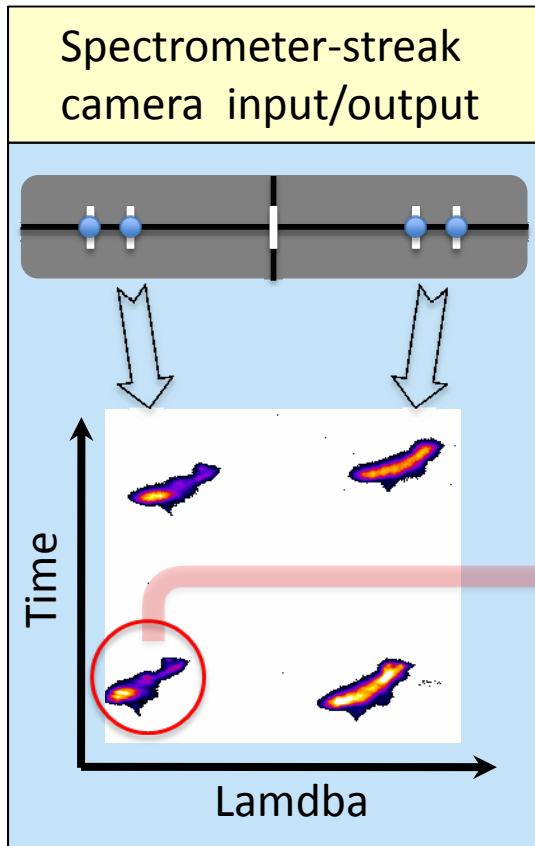
## NBI movie for a hohlraum with smaller capsule



Shot N091114

22 % SRS

# The 4 beams are multiplexed in the spectrometer-streak camera



- The SRS is  $\sim 20\%$  with AM  $\sim \pm 12\%$
- The pump AM is  $\sim \pm 4\%$
- A simple Tang growth model shows that these results are consistent with a *somewhat saturated* SRS

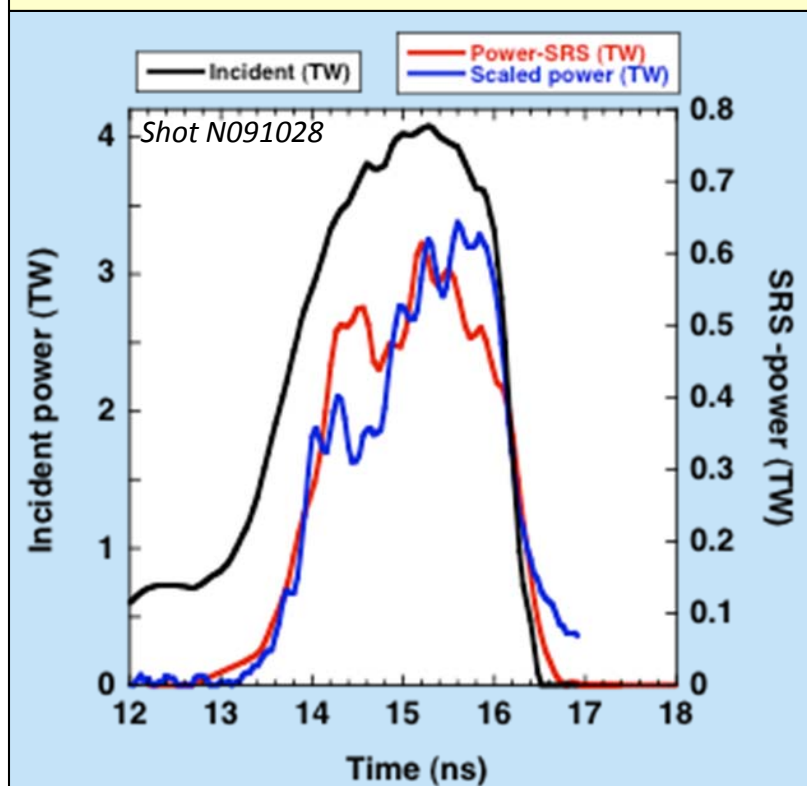


# Summary

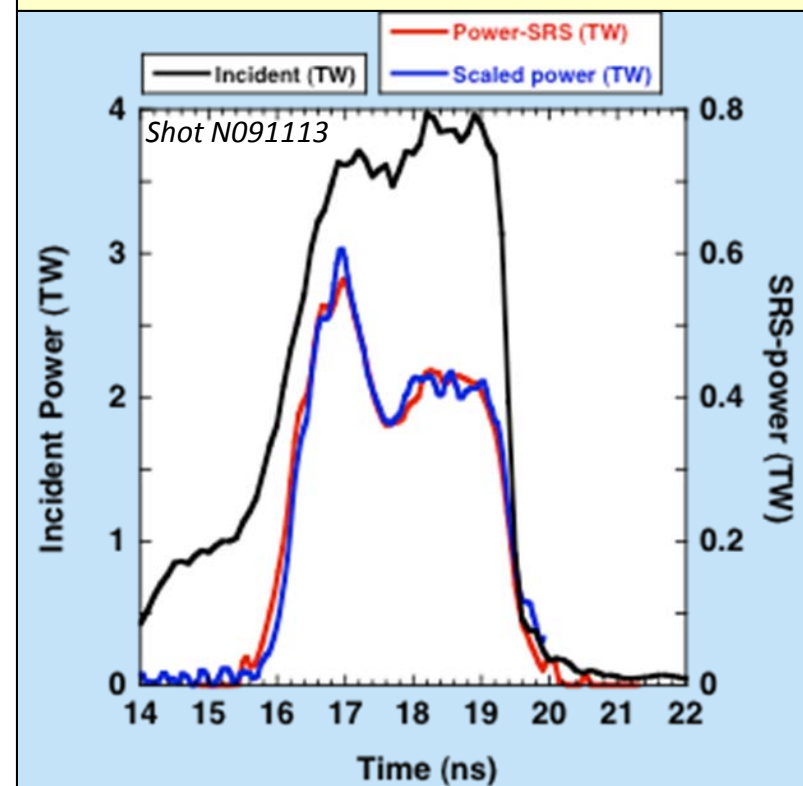
- Backscatter measurements have been activated on two quads in two beam cones on NIF
- The backscatter instrument consist of a FABS and NBI system which measure SRS and SBS separately
- Calibration techniques allow the instrument to achieve  $\sim 15\%$  error on the power measurement
- A new time-resolved NBI shows that the time-dependence of the entire backscatter signal is well represented by the FABS fast diode
- The backscatter measurements have guided decisions relating to the laser operation and the target design
- Future efforts will expand the NBI plate operation to include a quad at  $23.5^\circ$  from the hohlraum axis

# The NBI-TR shows a time-history that is similar to the FABS fast diode

720 kJ scale 0.9

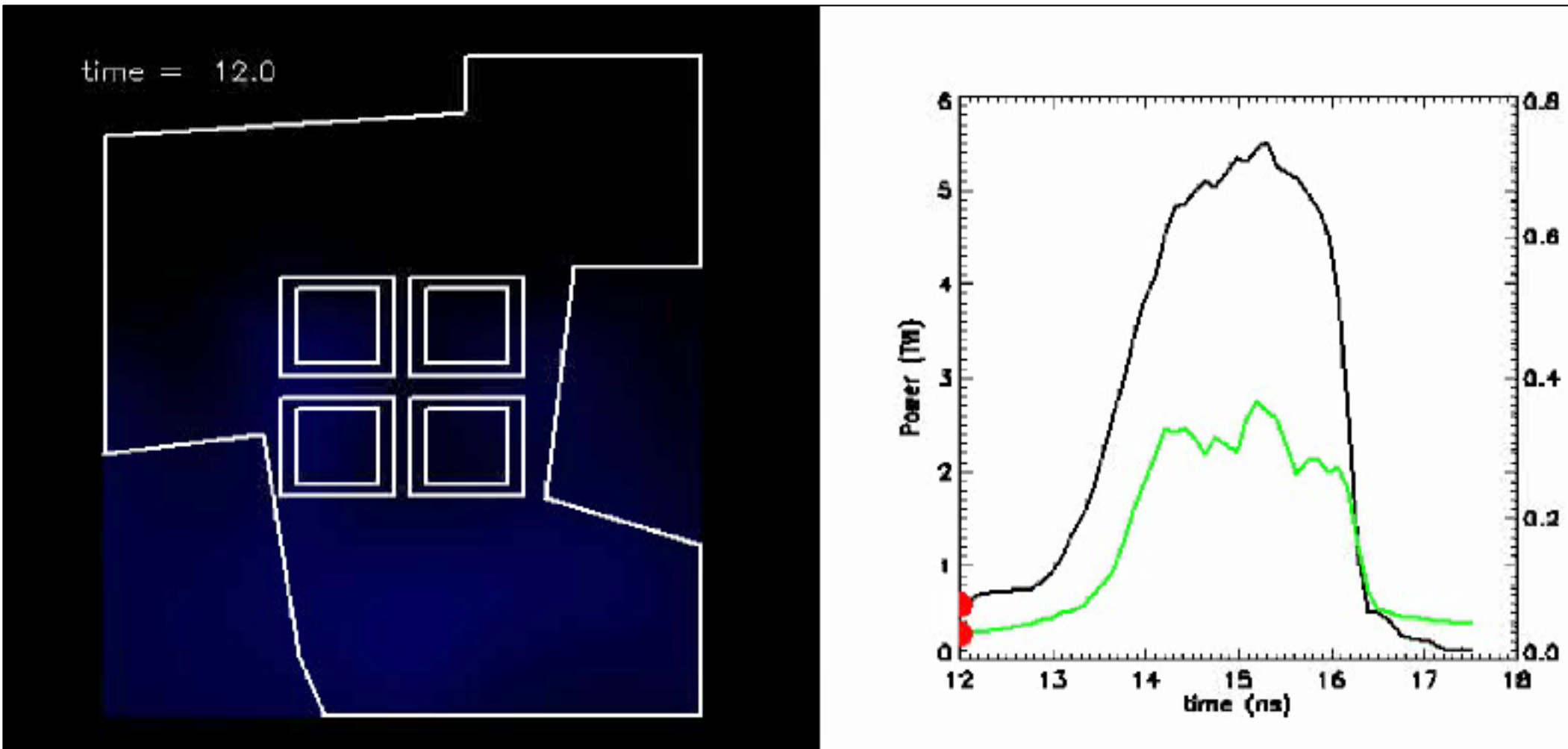


840 kJ scale 1.07



# End

## NBI movie showing significant AM

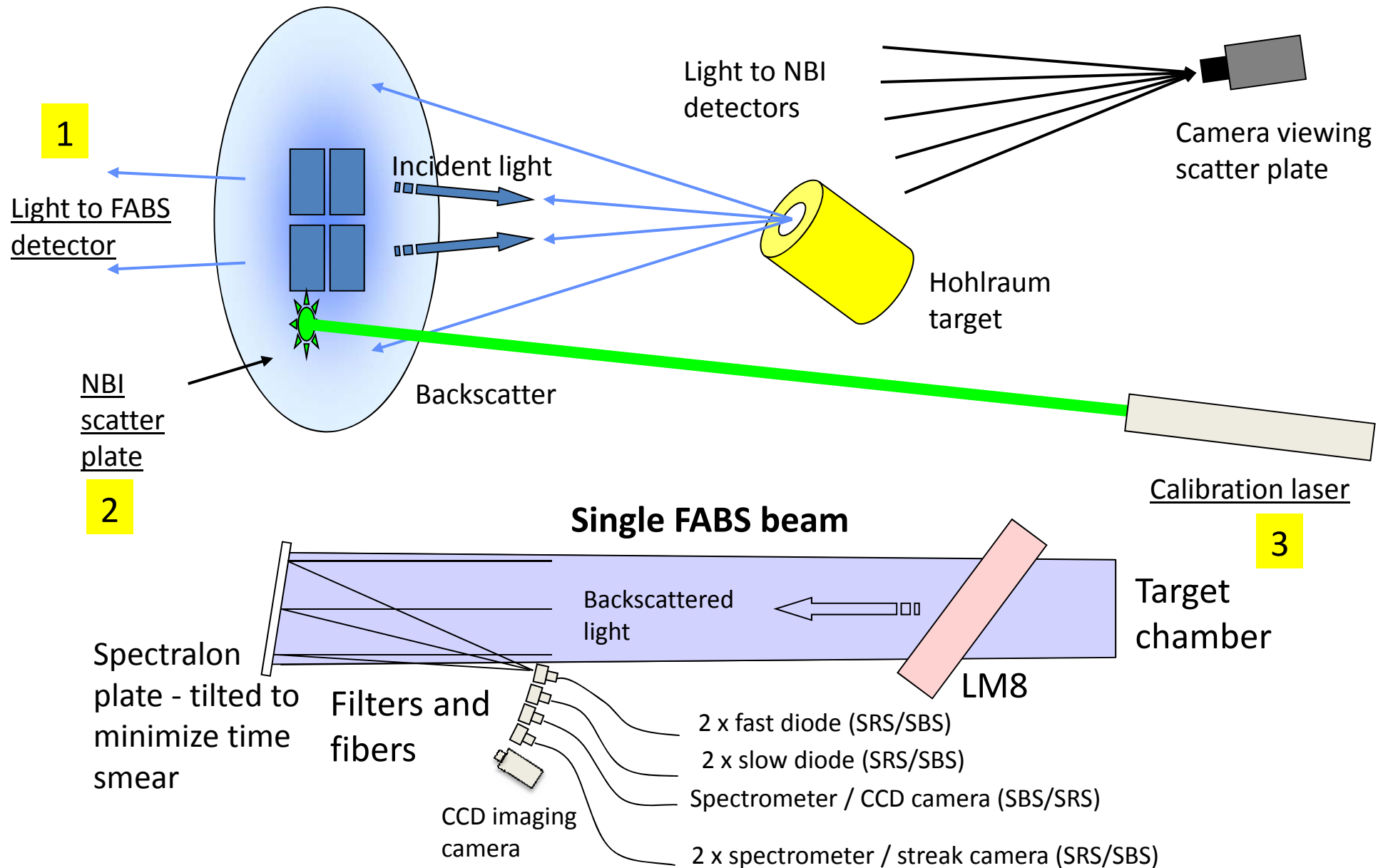


Shot N091025

24 % SRS



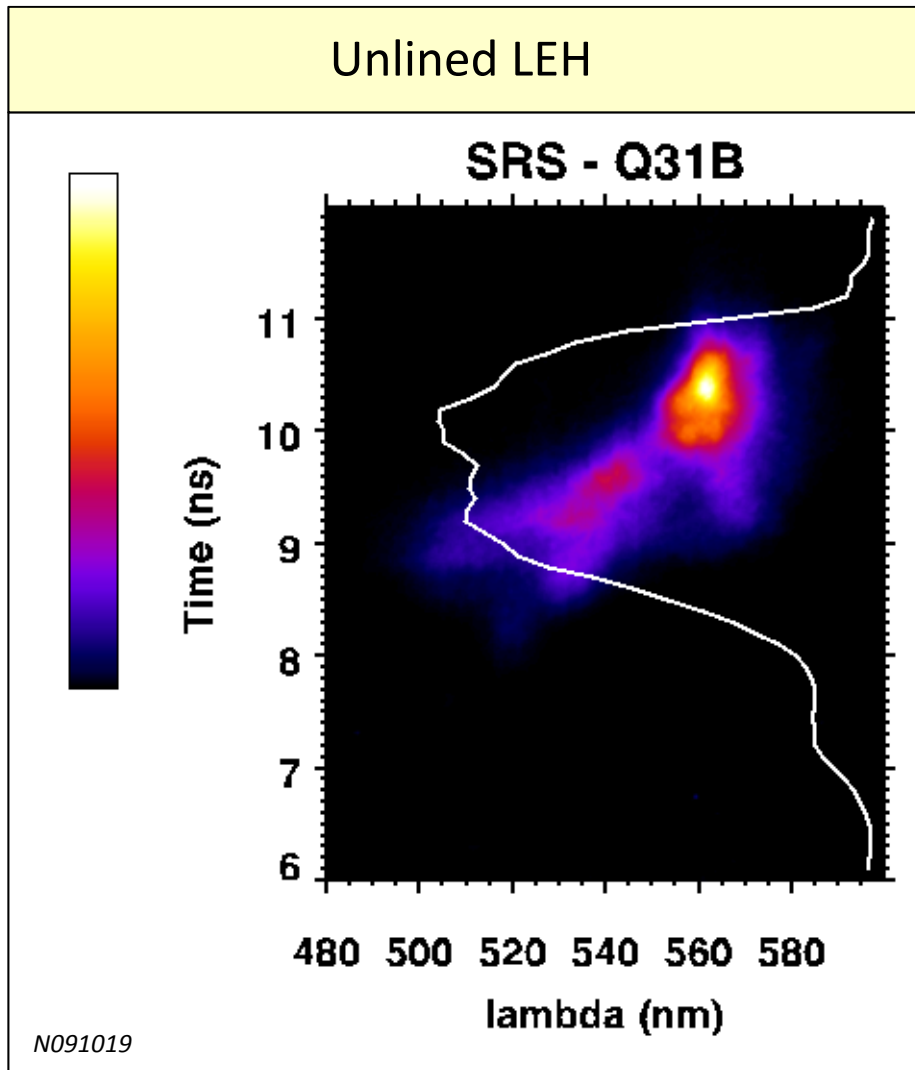
# The NIF system utilizes two instruments to measure backscatter light and one calibration instrument



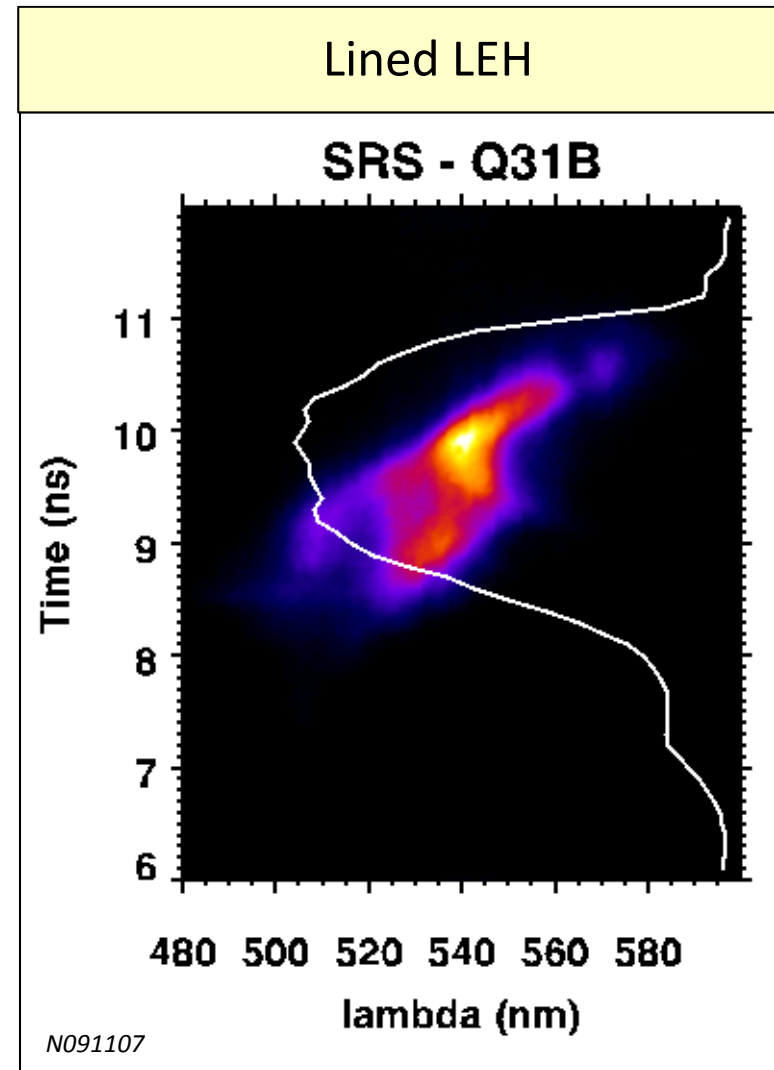
# SRS spectra shows laser propagation to higher density/temperature from unlined hohlraum experiments

LEH liners

NIC



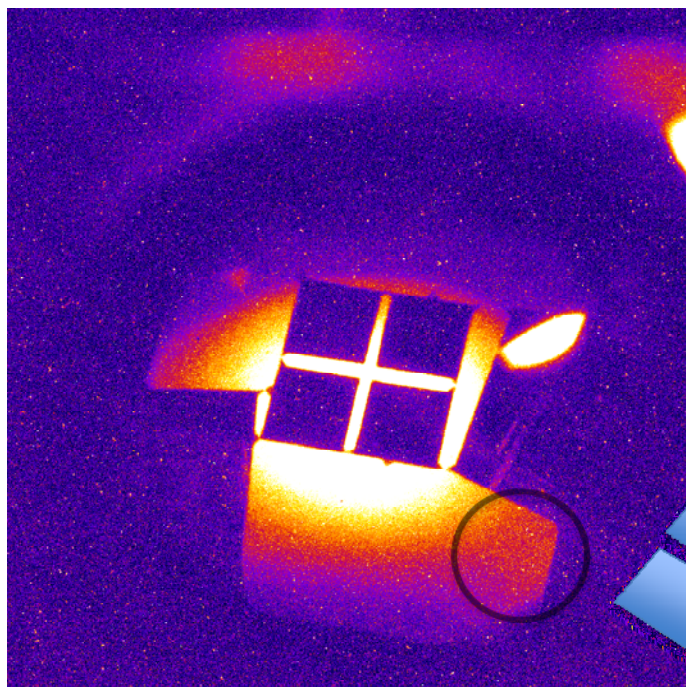
13 % SRS



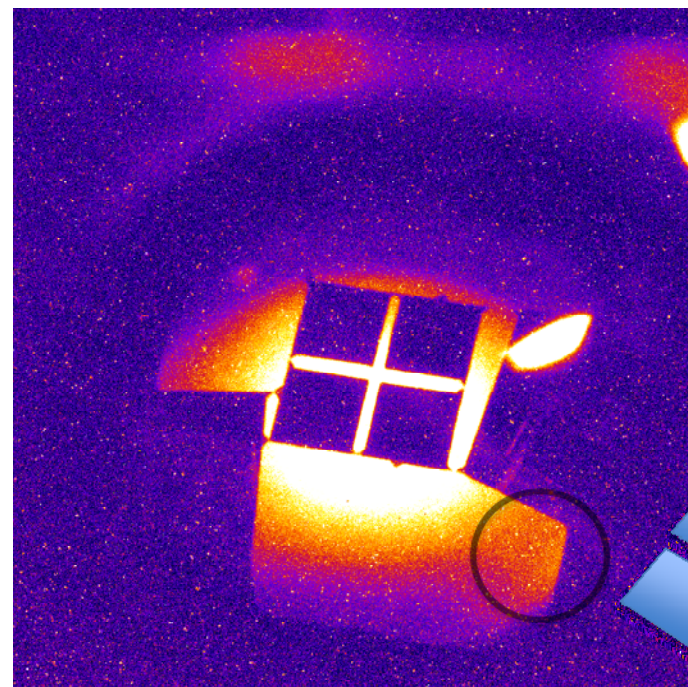
29 % SRS

## Future plans: measure SRS on 23.5 cone quad

N091030 ( $\Delta\lambda=3 \text{ \AA}$ )



N091030 ( $\Delta\lambda=8.5 \text{ \AA}$ )

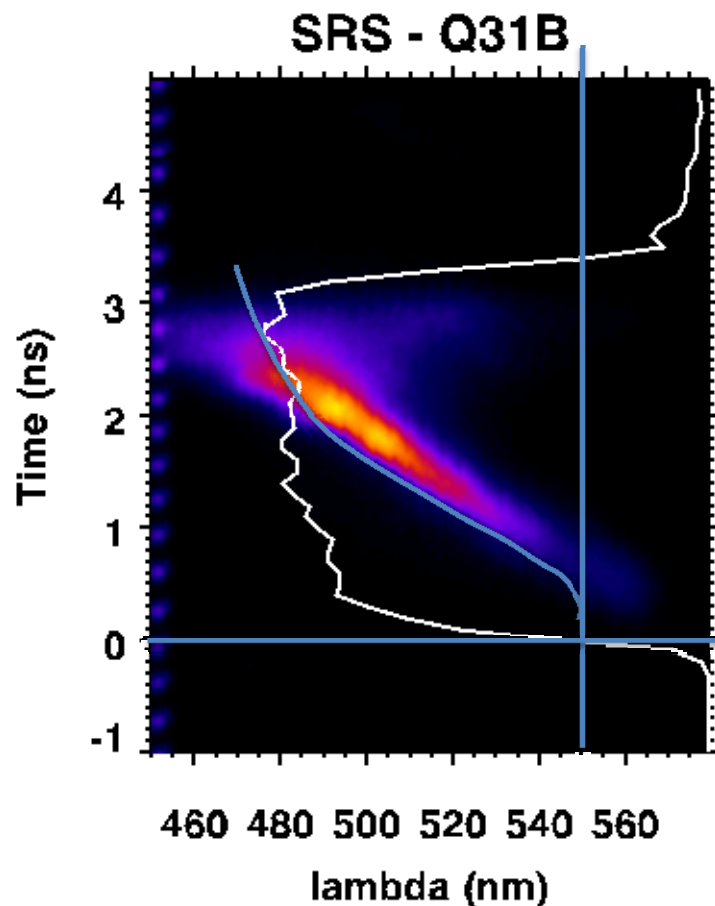


Approximate 23.5°  
quad location

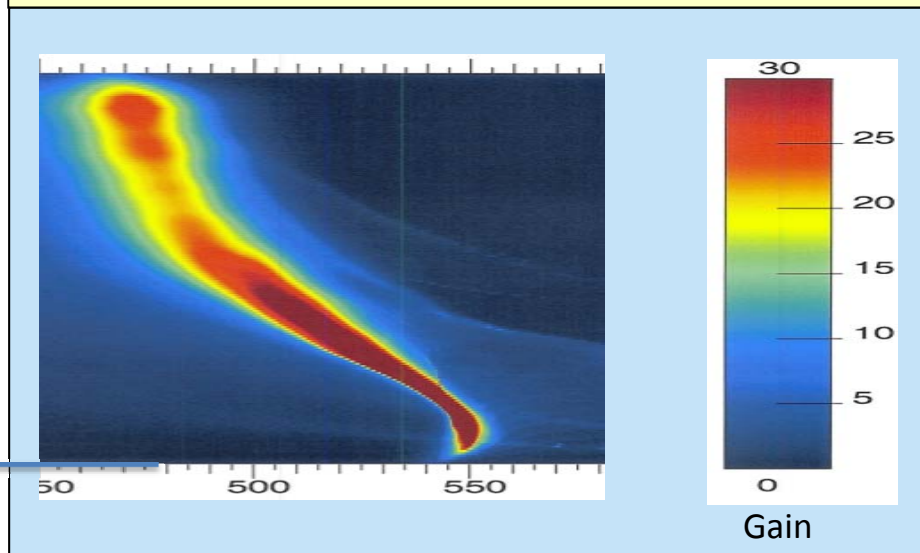
These shots and others show an increase in the light at the lower corner of the scatter plate

# SRS spectra from a C<sub>5</sub>H<sub>12</sub> gaspipe experiment is approximately reproduced in simulations

Stimulated Raman Scattering (SRS)  
spectral time-dependence data



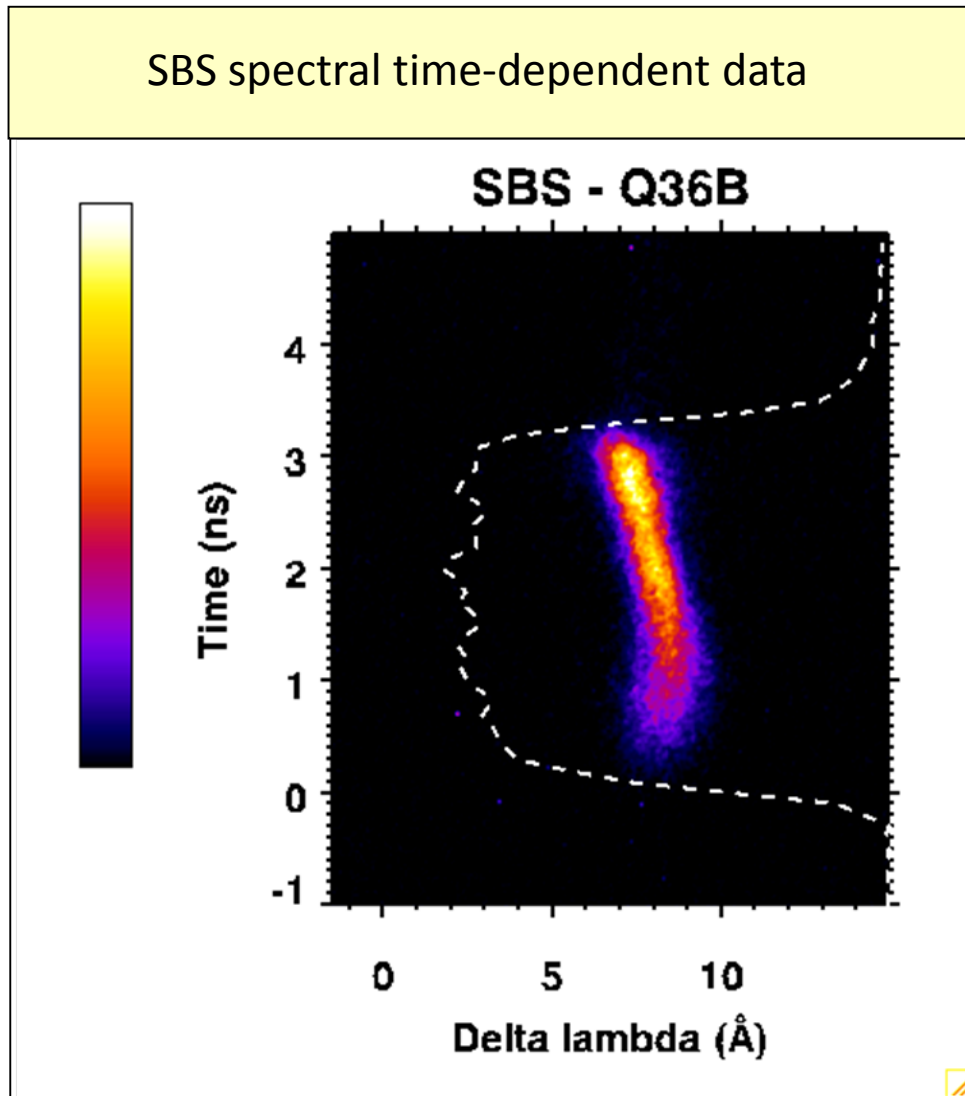
Stimulated Raman Scattering (SRS)  
spectral time-dependence simulation



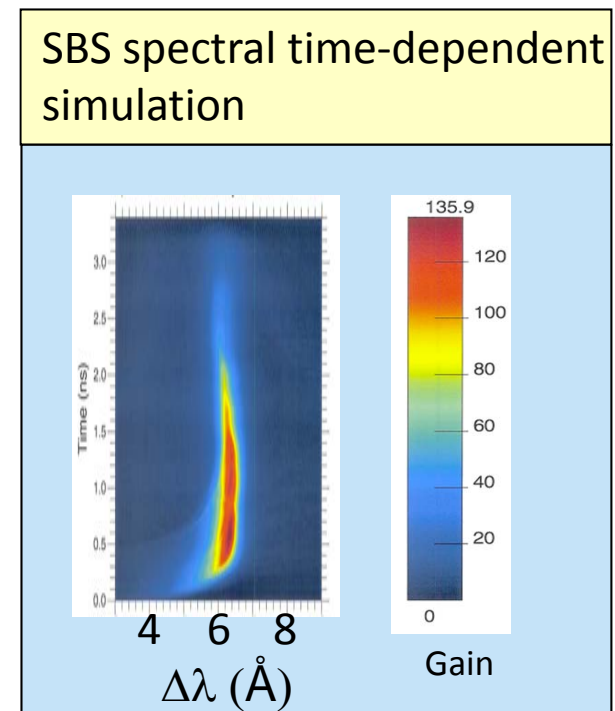
18 % SRS



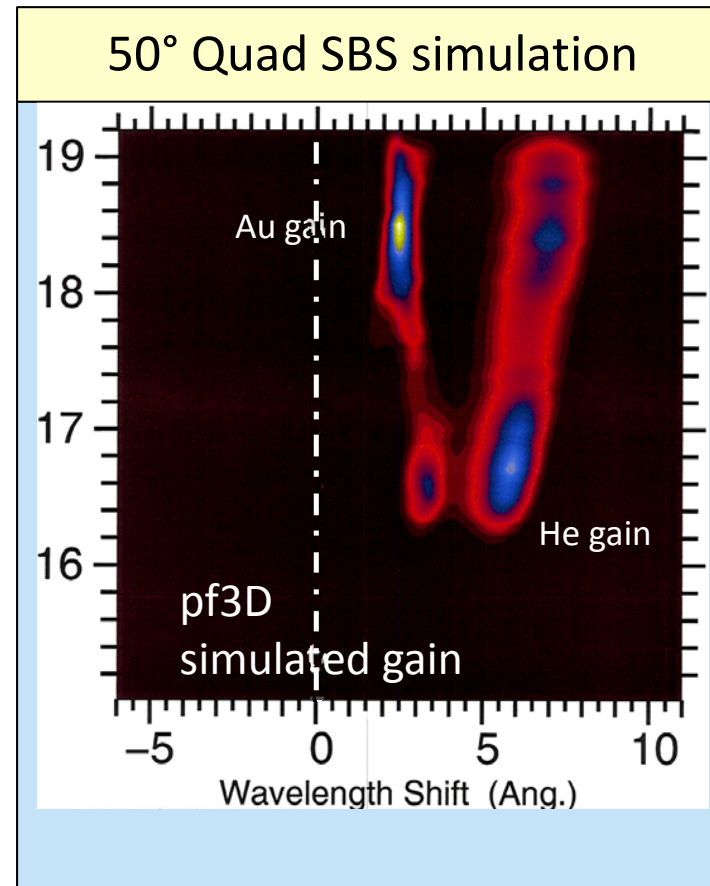
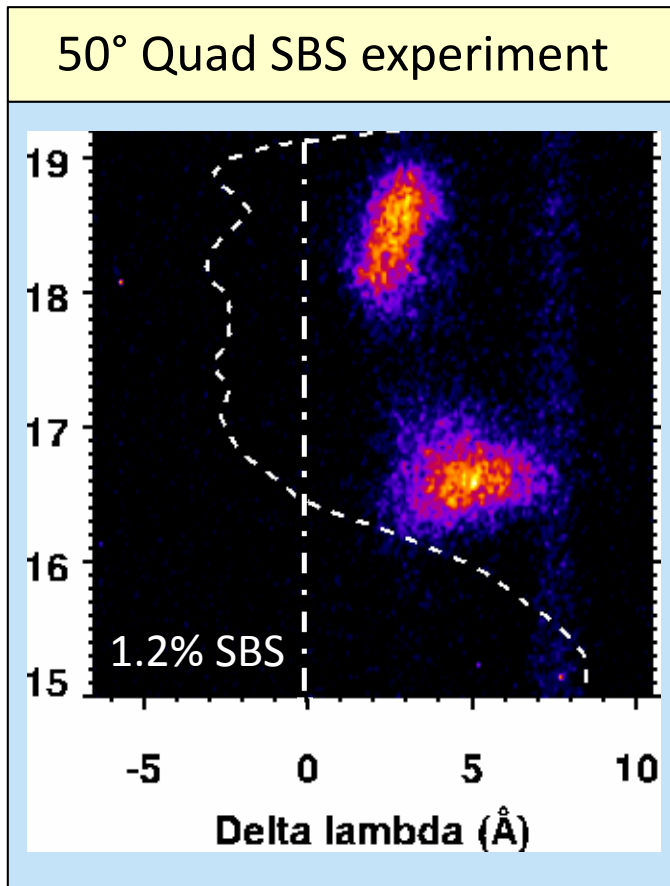
# SBS spectra from a CO<sub>2</sub> gaspipe experiment is approximately reproduced in simulations



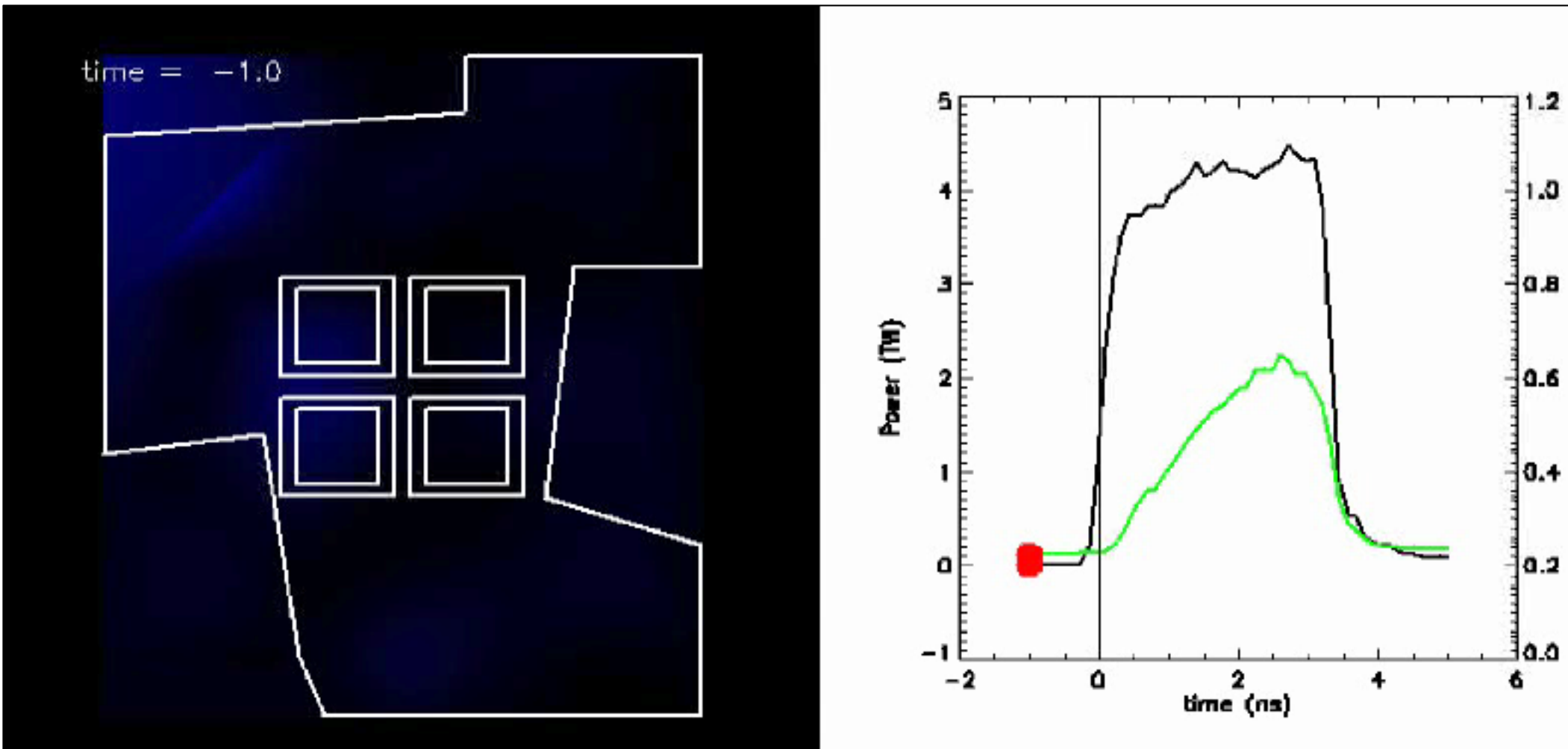
40 % SBS



# SBS outer beam spectra show good agreement with simulations



# NBI movie for the gaspipe

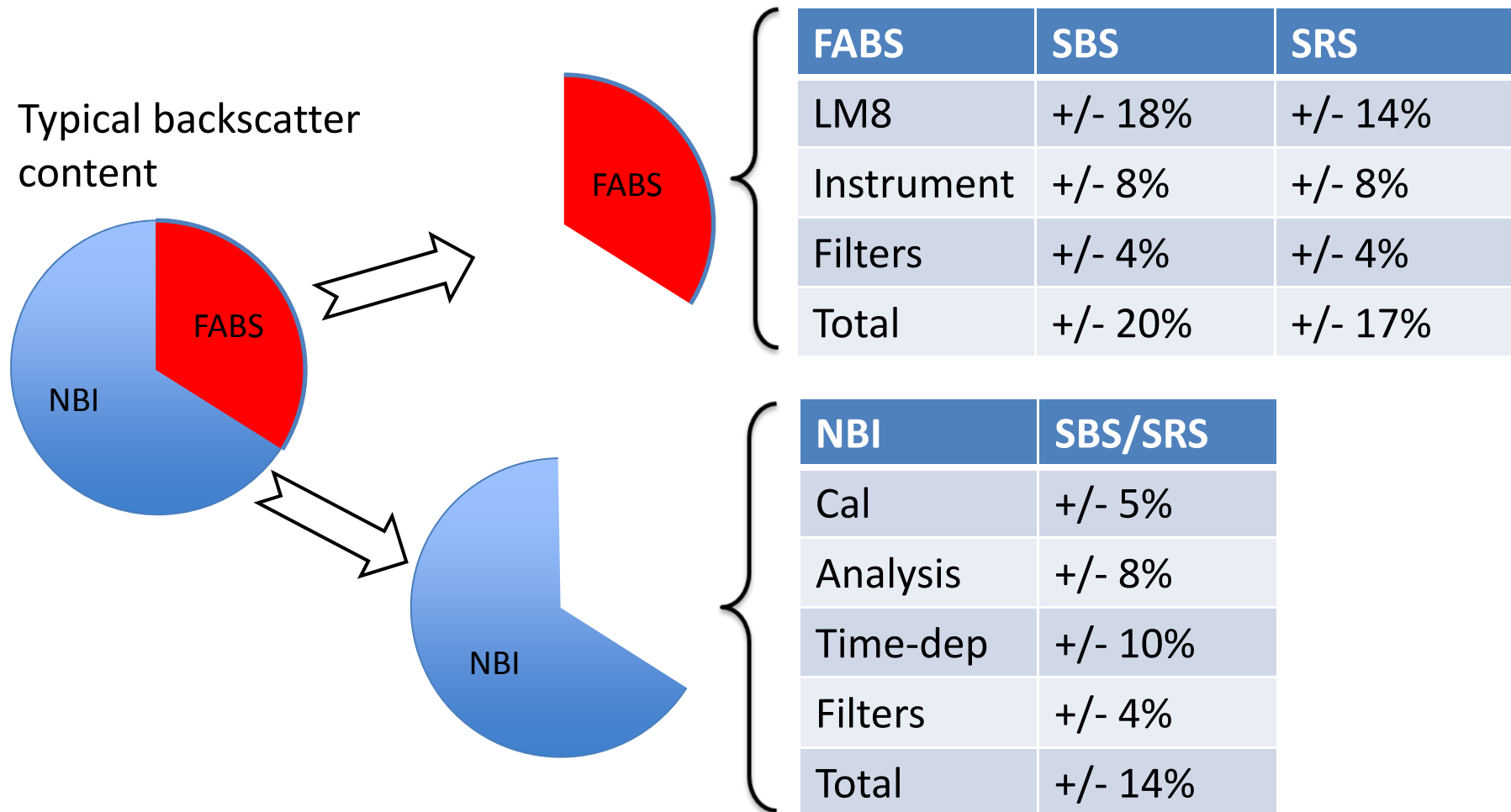


Shot N091202

19 % SRS

# Error summary

Typical backscatter content



$$\text{Overall error} = 20\% * 1/3 + 14\% * 2/3 = 16\%$$